

# Kinetics of electronically excited O<sub>2</sub> molecules in the mixture of CO<sub>2</sub>, CO, N<sub>2</sub>, O<sub>2</sub> gases

Kirillov A.S.<sup>1</sup>, Werner R.<sup>2</sup>, Guineva V.<sup>2</sup>



<sup>1</sup>Space Research and Technology Institute, Stara Zagora Department, BAS, Bulgaria

<sup>2</sup>Polar Geophysical Institute (PGI), PAH, Apatity, Murmansk region, Russia



# Introduction (1)

CO<sub>2</sub>, CO, N<sub>2</sub>, O<sub>2</sub> gases are the main components in the atmospheres of terrestrial planets. Electronically excited molecules play very important role in chemical kinetics of a mixture of atmospheric gases. It has been proposed by Toumi [1993], Siskind et al. [1993], Prasad [1997] that the interaction of O<sub>2</sub>(b<sup>1</sup>Σ<sub>g</sub><sup>+</sup>) with H<sub>2</sub> and N<sub>2</sub>O molecules may significantly influence the chemistry of the stratosphere and upper troposphere of the Earth as a source of odd-hydrogen (HO<sub>x</sub>) or odd-nitrogen (NO<sub>x</sub>).

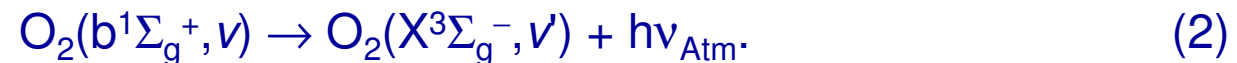
O<sub>2</sub>(b<sup>1</sup>Σ<sub>g</sub><sup>+</sup>) molecules are effectively produced in the atmosphere of the Earth at altitudes between 80 and 110 km. Three-body collisions



are main production mechanism of electronically excited molecular oxygen at the altitudes of the nightglow in the atmosphere of the Earth [Slanger and Copeland, 2003].

## Introduction (2)

Spontaneous radiative transitions from the  $b^1\Sigma_g^+$  state to the ground  $X^3\Sigma_g^-$  state lead to emissions of Atmospheric bands in the nightglow of the atmosphere.



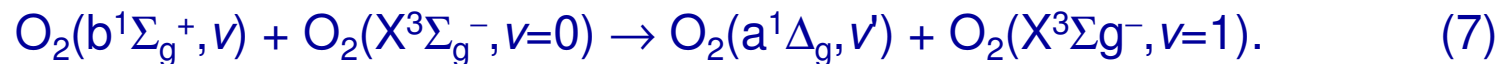
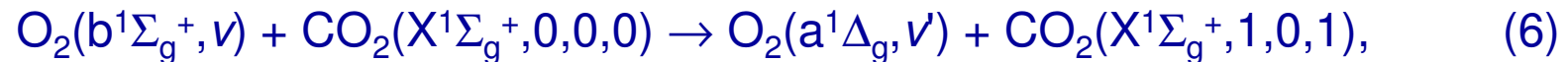
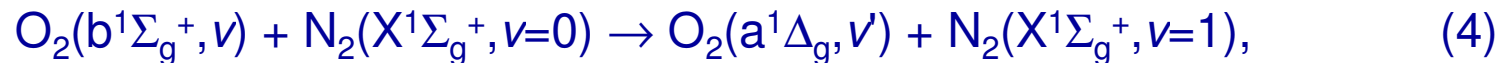
A study of excitation, quenching, and energy transfer processes in the oxygen nightglow on Venus and Mars by Krasnopolsky [2011] has considered altitude profiles of  $[\text{O}_2(b^1\Sigma_g^+)]$  in atmospheres of the planets. The study was based on the observed nightglow intensities and vertical profiles, measured reaction rate coefficients, and photochemical models of the nighttime atmospheres of the Venus and Mars. Krasnopolsky [2011] has analysed mainly the kinetics of  $\text{O}_2(b^1\Sigma_g^+, \nu=0)$ , but detailed study of  $\text{O}_2(b^1\Sigma_g^+, \nu \geq 0)$  needs a set of the quenching constants for collisions with molecules of main gases in atmospheres of the planets.

# The quenching rate coefficients (1)

We apply the Rosen-Zener approximation to calculate the removal rates of  $O_2^*$  in inelastic collisions with  $CO_2$ ,  $CO$ ,  $N_2$ ,  $O_2$  molecules [Kirillov, 2004a,b]. The quenching rates of the  $b^1\Sigma_g^+, v=1-15$  state by  $O_2$  molecules are calculated with the consideration of intermolecular (EE-processes) transfers of electronic excitation [Kirillov, 2012]



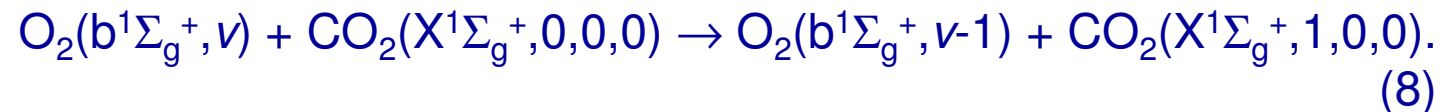
Also we consider EV-processes of removal in the collisions with  $N_2$ ,  $CO$ ,  $CO_2$ ,  $O_2$  molecules



Here there are the transitions of electronically excited  $O_2^*$  molecules in another excited state with vibrational excitation of the ground state of target molecules.

## The quenching rate coefficients (2)

Moreover, we take into account the processes of VV-processes of energy exchange in the collisions with carbon dioxide molecules

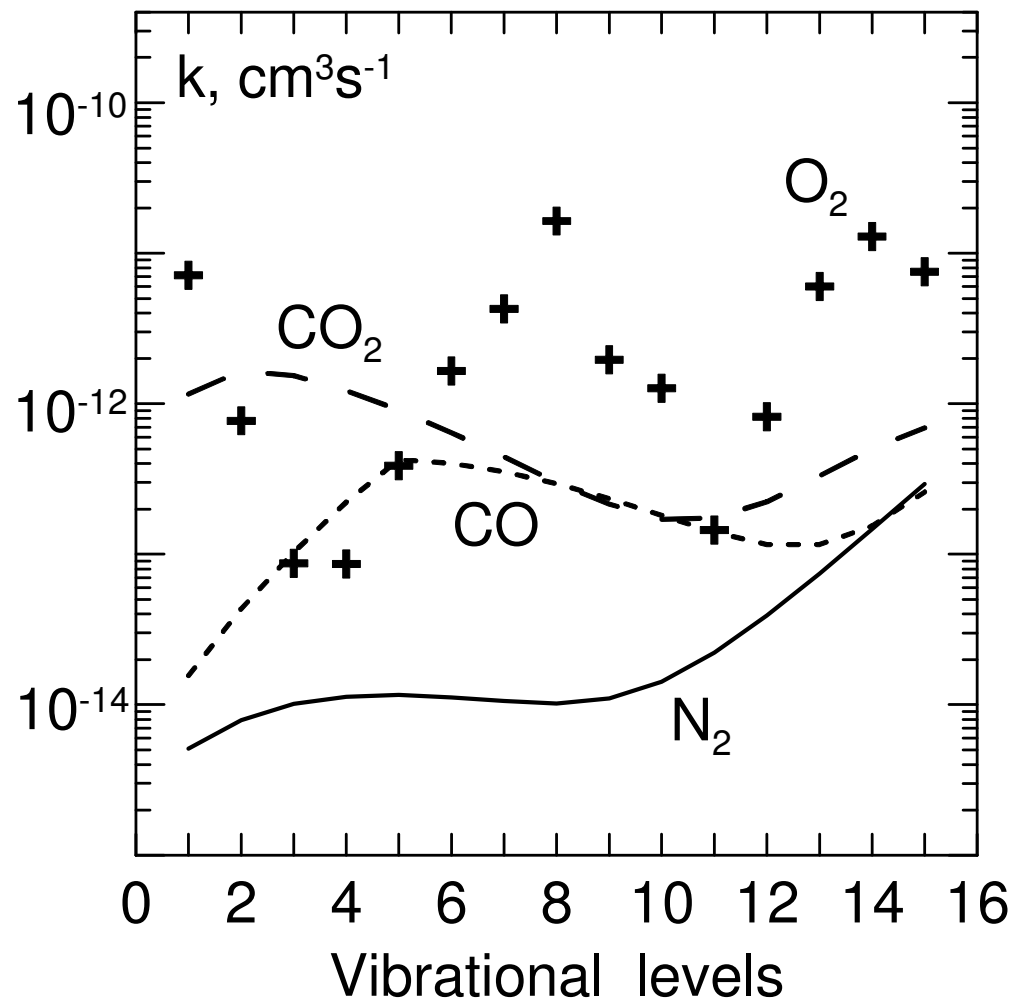


The process (8) means that the  $\text{O}_2(\text{b}^1\Sigma_g^+, \nu)$  molecule loses one vibrational quantum of the  $\text{b}^1\Sigma_g^+$  state and the energy transforms in vibrational quantum of symmetric stretch mode of  $\text{CO}_2$  molecule.

The calculated quenching rate coefficients for the  $\text{b}^1\Sigma_g^+, \nu=1-15$  state in the collisions (3-8) with  $\text{O}_2$ ,  $\text{N}_2$ ,  $\text{CO}$ ,  $\text{CO}_2$  molecules are presented in Fig.1. The calculations have shown that the quenching is very efficient in the collisions with  $\text{O}_2$  molecules for high vibrational levels of this singlet state.

More effective quenching by carbon dioxide molecules is seen for lower vibrational states of the  $\text{b}^1\Sigma_g^+$  state. The calculated quenching rate coefficients of electronically excited  $\text{O}_2$  molecules are used in the simulations of vibrational populations of  $\text{O}_2$  electronic states in the mixture of gases.

# The quenching rate coefficients (3)



**Fig.1.** The calculated quenching rate coefficients for the  $b^1\Sigma_g^+, v=1-15$  state in the collisions with  $\text{O}_2$ ,  $\text{N}_2$ ,  $\text{CO}$ ,  $\text{CO}_2$  molecules (crosses; solid, short-dashed, long-dashed lines, respectively).

# Vibrational populations of $O_2(b^1\Sigma_g^+)$ in the mixture (1)

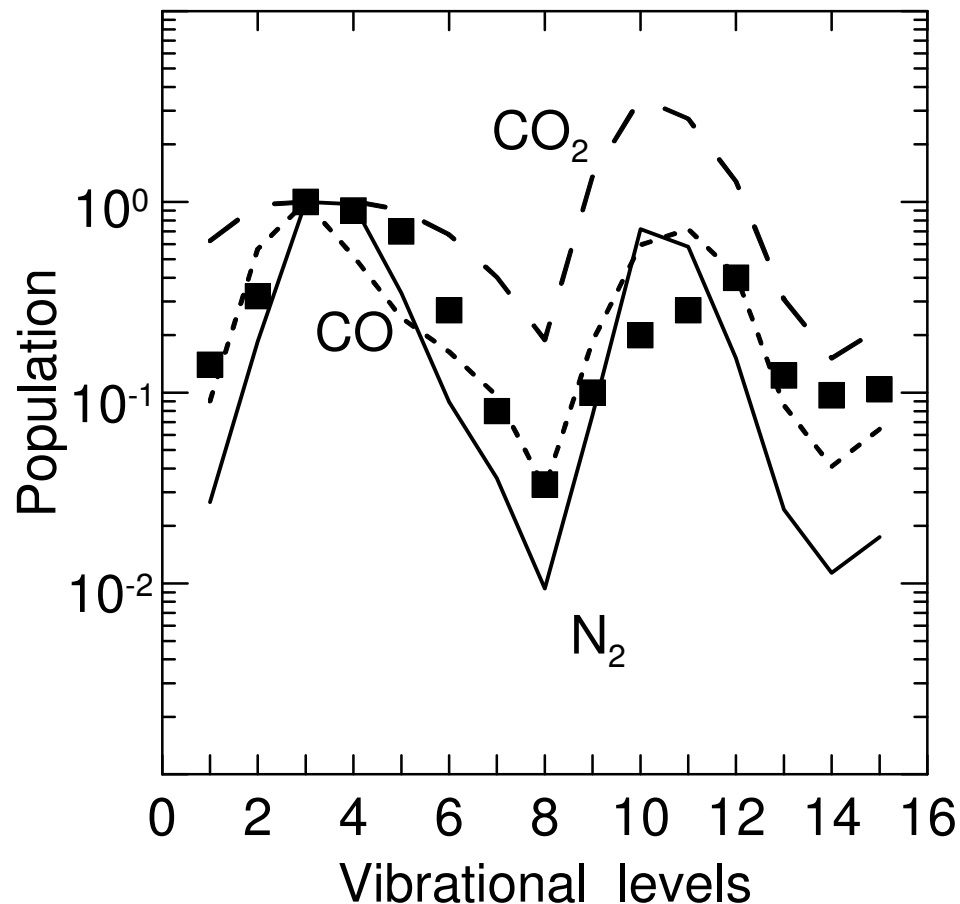
It is suggested that three-body collisions (1) are the production mechanism of initially excited  $O_2$ . Vibrational populations of singlet oxygen are calculated at mixture pressures of 10-1-100 Pa. The removal rates are taken according to the data presented in Fig.1.

The calculated vibrational populations of singlet  $b^1\Sigma_g^+, v=1-15$  state of molecular oxygen in the mixture of  $O_2^*$  with the gases  $N_2$ ,  $CO$ ,  $CO_2$  at pressures of 10-1-100 Pa are shown in Fig.2. Also the experimental data of populations estimated from spectrometric observations from Keck-I telescope [Slanger et al., 2000] are presented in Fig.2. Slanger et al. [2000] have observed the intensities of Atmospheric bands (2) emitted from fifteen vibrational levels of the singlet state in the nightglow of the Earth.

It is seen from this figure that the behaviour of the populations in the cases of  $N_2-O_2$  and  $CO-O_2$  mixtures are similar. The principal influence of carbon dioxide molecules is seen in the case of  $CO_2-O_2$  mixture. This mixture is an analogue of Venus and Mars atmosphere at the altitudes of the nightglow. It is seen from Fig.2 that there is an increase of relative populations of highest vibrational levels in comparison with ones of lowest levels. The difference in the behaviour of the populations for the mixtures can be explained by the removal rates presented in Fig.1.

## Vibrational populations of $O_2(b^1\Sigma_g^+)$ in the mixture (2)

The results of the calculations can be applied in the study of electronic kinetics of  $O_2^*$  at the altitudes of nightglows in the atmospheres of terrestrial planets and in active mediums of laboratory discharges and lasers.



**Fig.2.** The calculated vibrational populations of singlet  $b^1\Sigma_g^+, v=1-15$  state of molecular oxygen in the mixture of  $O_2^*$  with the gases  $N_2$  (solid line),  $CO$  (short-dashed line),  $CO_2$  (long-dashed line); experimental data by Slinger et al . [2000] – squares.



# Conclusions

**The main results of this presentation are as follows:**

1. Applying the Rosen-Zener approximation we have calculated the removal rates of  $O_2(b^1\Sigma_g^+, v)$  in inelastic collisions with  $CO_2$ ,  $CO$ ,  $N_2$ ,  $O_2$  molecules. The calculation has taken into account EE (3), EV (4-7) and VV (8) electron energy transfer processes.
2. Using the calculated removal rate constants we have simulated vibrational populations of singlet  $b^1\Sigma_g^+, v=1-15$  state of molecular oxygen in the mixture of  $O_2^*$  with the gases  $N_2$ ,  $CO$ ,  $CO_2$  at pressures of 10-1-100 Pa. The results of the calculation show the principal dependence of the populations on the kind of admixture gas. The dependence can be explained by peculiarities of removal rates of  $O_2(b^1\Sigma_g^+)$  in inelastic collisions with another molecules.

# References

- Kirillov A.S. Application of Landau-Zener and Rosen-Zener approximations to calculate rates of electron energy transfer processes. // *Adv. Space Res.*, 2004a, v.33, p.993–997.
- Kirillov A.S. Calculation of rate coefficients of electron energy transfer processes for molecular nitrogen and molecular oxygen. // *Adv. Space Res.*, 2004b, v.33, p.998–1004.
- Kirillov A.S. Calculation of rate coefficients for the interaction of singlet and triplet vibrationally excited oxygen. // *Quantum Electronics*, 2012, v.42, p.653-658.
- Krasnopolsky V.A. Excitation of the oxygen nightglow on the terrestrial planets. // *Planet. Space Sci.*, 2011, v.59, p.754-766.
- Prasad S.S. Potential atmospheric sources and sinks of nitrous oxide. 2. Possibilities from excited  $O_2$ , "embryonic"  $O_3$ , and optically pumped excited  $O_3$ . // *J. Geophys. Res.*, 1997, v.102, p.21527-21536.
- Siskind D.E., Summers M.E., Mlynczak M.G. An evaluation of  $O_2(b^1\Sigma_g)$  as a possible source of OH and odd-nitrogen in the stratosphere and mesosphere. // *Geophys. Res. Lett.*, 1993, v.20, p.2047-2050.
- Slanger T.G. and Copeland R.A. Energetic oxygen in the upper atmosphere and the laboratory. // *Chem. Rev.*, 2003, v.103, p.4731-4766.
- Slanger T.G., Cosby P.C., Huestis D.L., Osterbrock D.E. Vibrational level distribution of  $O_2(b^1\Sigma_g^+, v=0-15)$  in the mesosphere and lower thermosphere region. // *J. Geophys. Res.*, 2000, v.105, p.20557-20564.
- Toumi R. A potential new source of OH and odd-nitrogen in the atmosphere. // *Geophys. Res. Lett.*, 1993, v.20, p.25-28.

**Thank you  
very much  
for your attention**