

- **Geophysically Induced Currents, a space weather hazard.**
- **Case study – Europe under intense geomagnetic storms of the solar cycle 23**

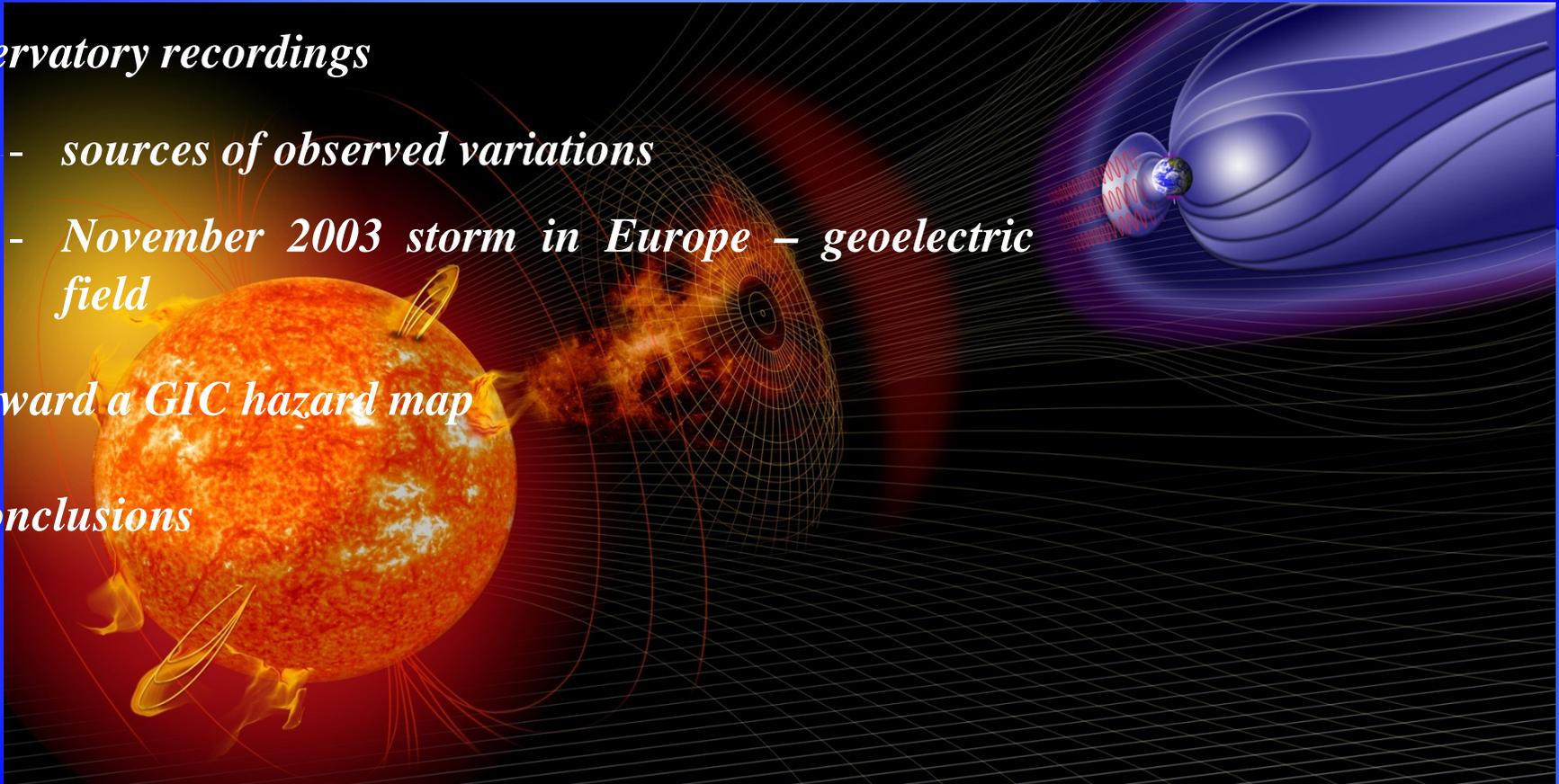
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Institute of Geodynamics, Bucharest, Romania, crisan@geodin.ro

Acknowledgements: Funding by the Ministry of Education and Research, UEFISCDI, Project IDEI 93/2011

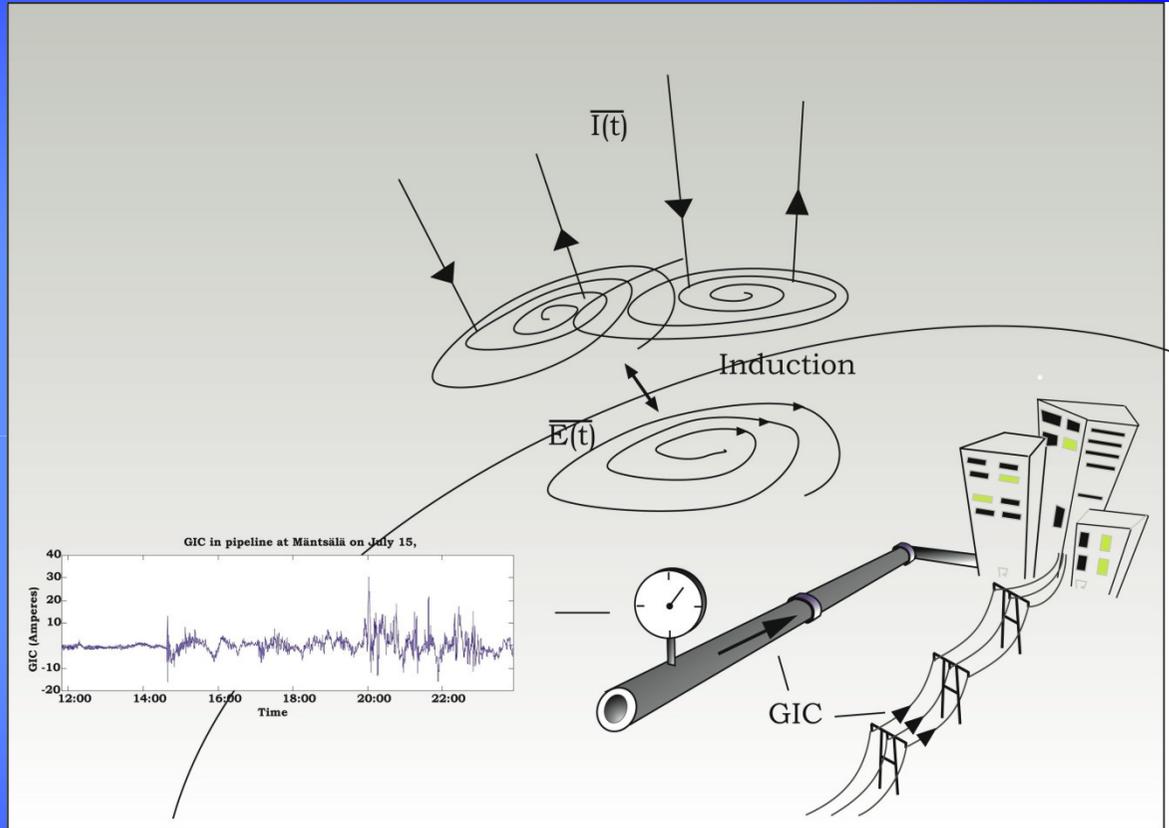
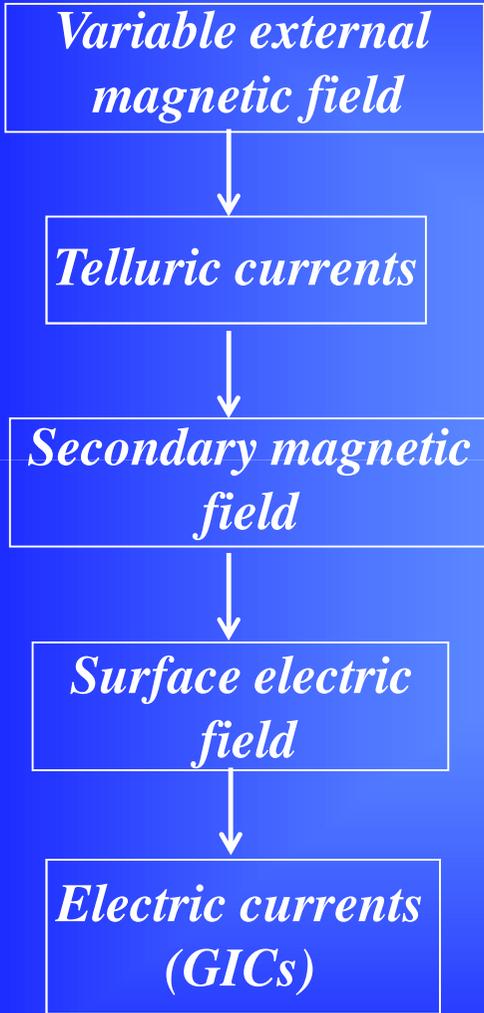
Outline

- *Motivation*
- *Intense storms ($Dst < -150$ nT) of cycle 23*
- *Calculation of geoelectric field from geomagnetic observatory recordings*

- *sources of observed variations*
- *November 2003 storm in Europe – geoelectric field*
- *Toward a GIC hazard map*
- *Conclusions*

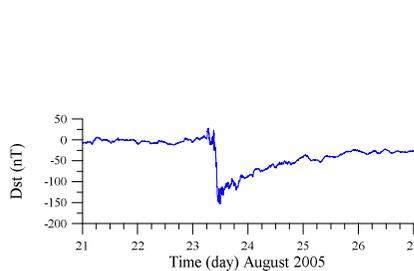
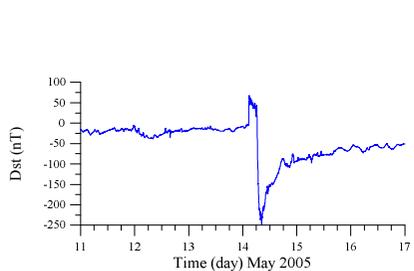
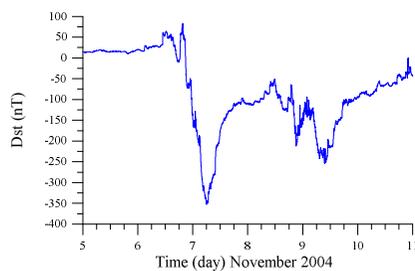
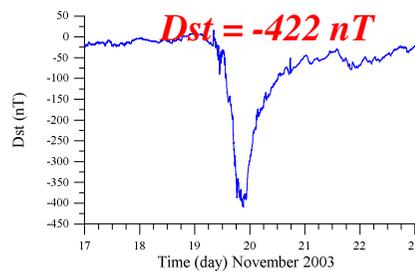
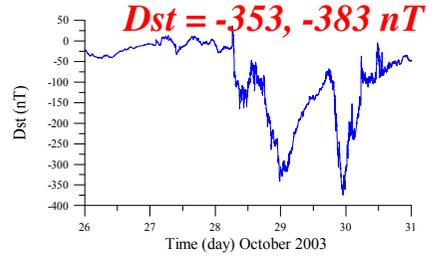
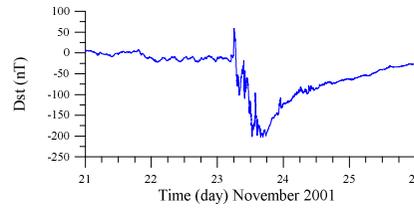
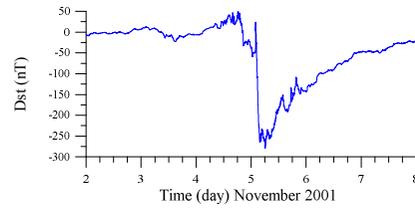
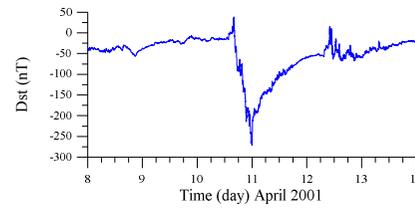
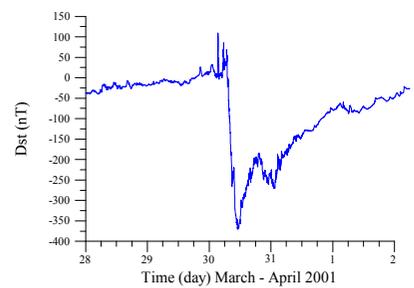
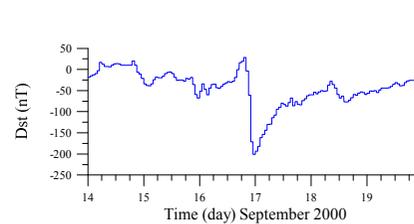
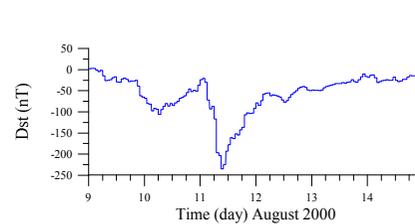
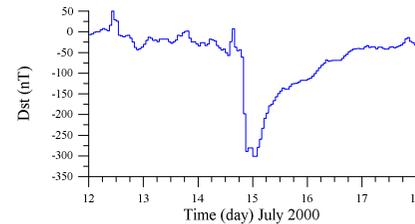
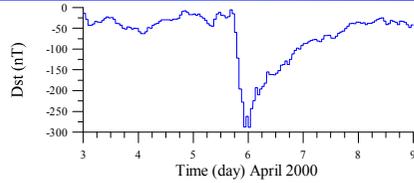
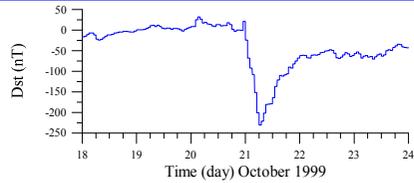
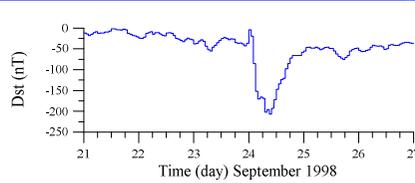
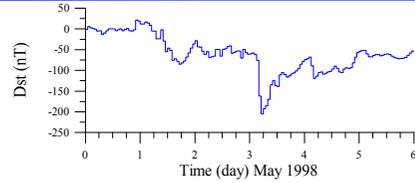


Motivation



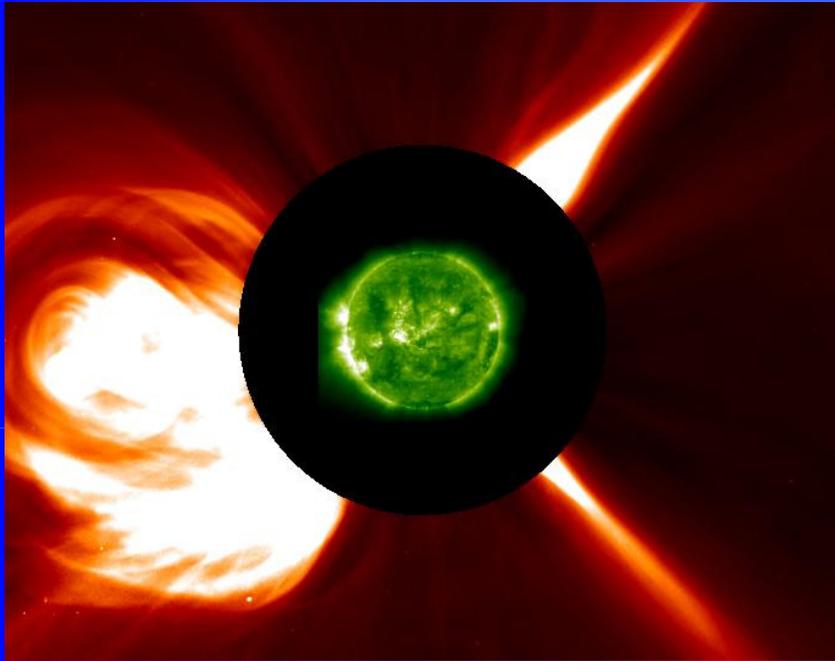
<http://en.wikipedia.org/>

Intense ($Dst < -150$ nT) storms – cycle 23



November 2003 storm

CME



November 18, 2003

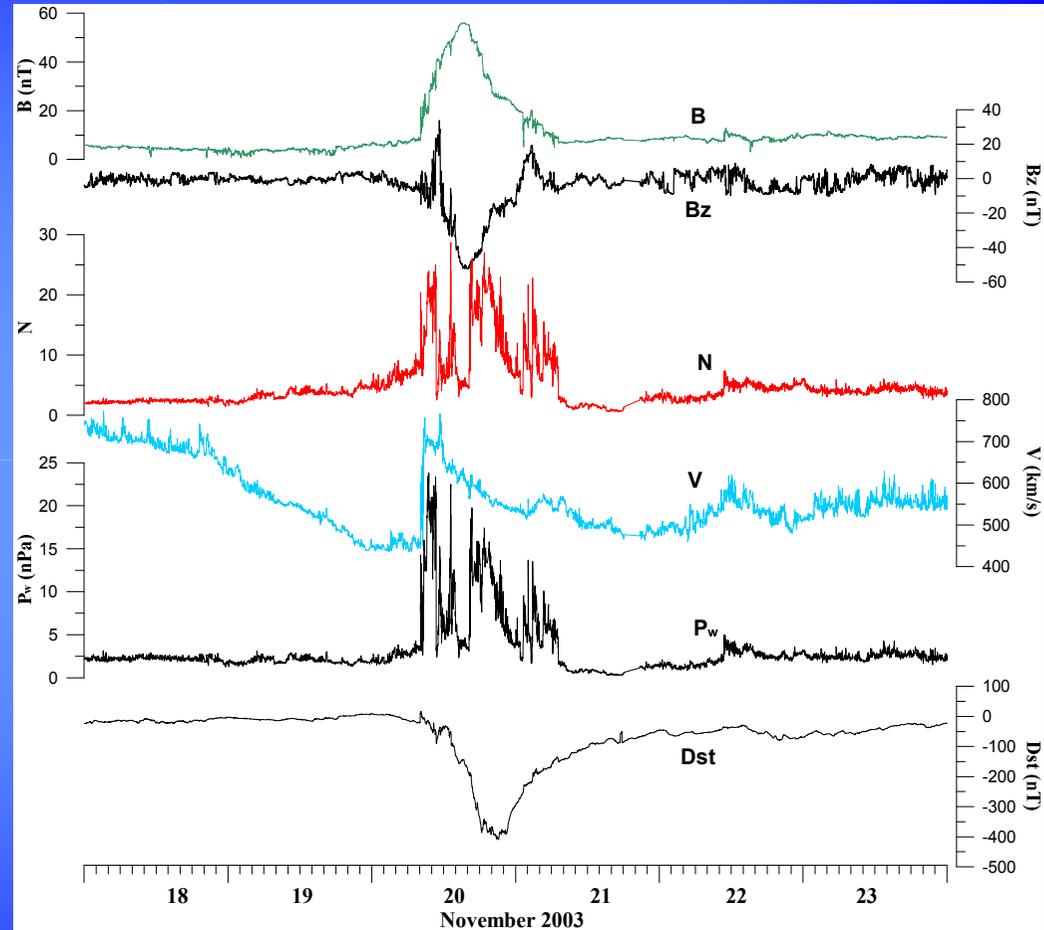
10:24 UT

GOES

LASCO C2 image (jhelioviewer.org/)

Solar eruption November 18, 2003, 8:12 UT

ICME



www.omniwebdata

SSC, November 20, 2003

8:03 UT

Surface geoelectric field (1)

$$E_x(\omega) = \frac{Z(\omega)}{\mu_0} B_y(\omega), E_y(\omega) = \frac{Z(\omega)}{\mu_0} B_x(\omega)$$

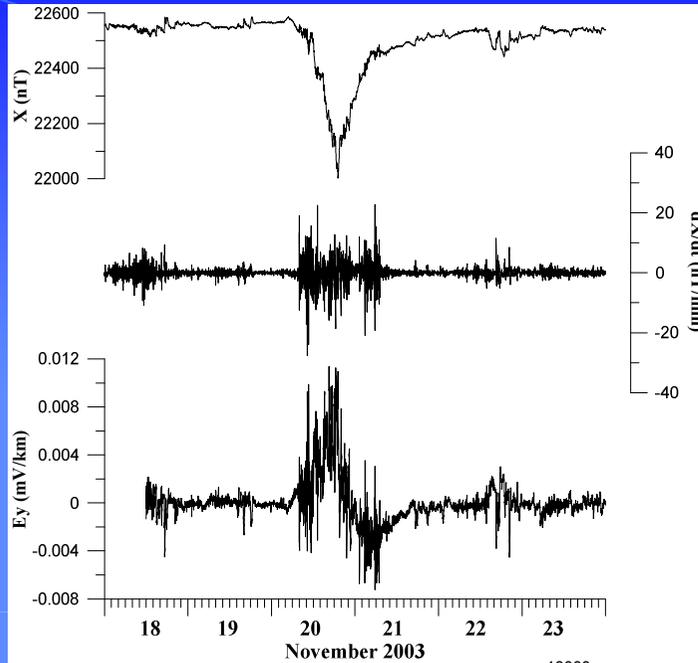
$$E_y(t) = -\frac{1}{\sqrt{\pi\mu_0\sigma}} \int_{-\infty}^t \frac{g_x(u)}{\sqrt{t-u}} du$$

$$E(T_N) = \frac{2}{\sqrt{\pi\mu_0\sigma}} (R_{N-1} - R_N - \sqrt{M} b_{N-M})$$

$$R_N = \sum_{n=N-M+1}^N b_n \sqrt{N-n+1}$$

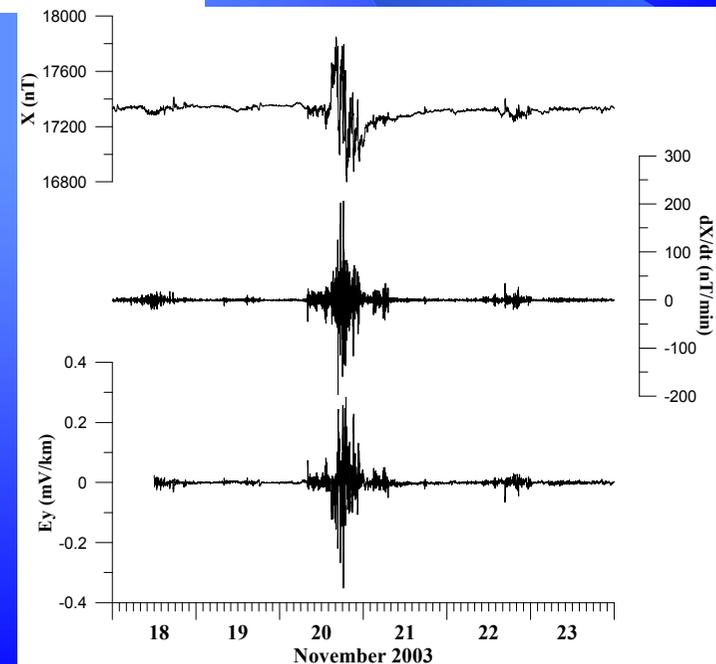
Viljanen & Pirjola, 1989

$$E(T_N) = \sqrt{E_x^2 + E_y^2}$$



SUA

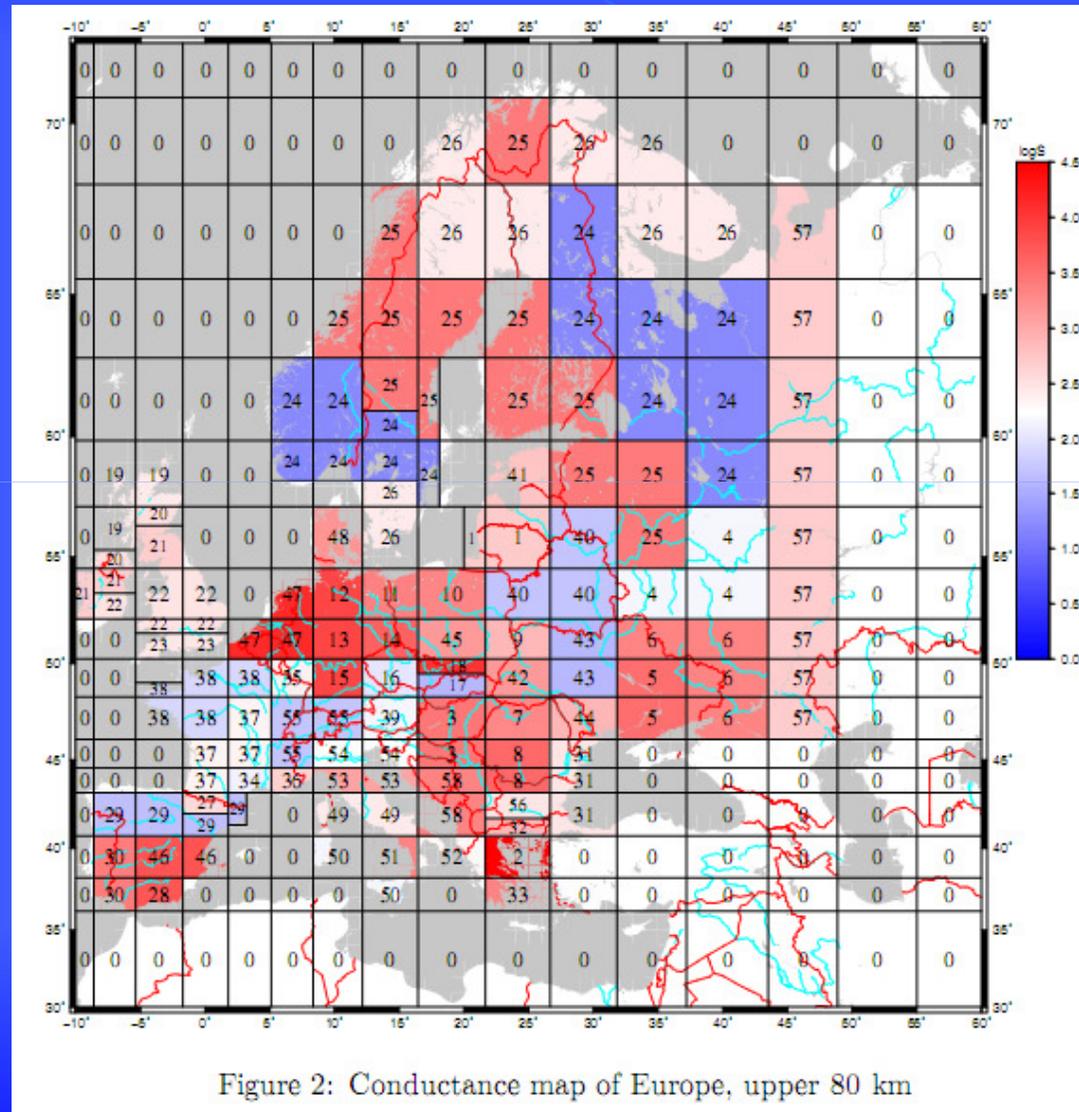
ESK



Intermagnet observatories

Surface geoelectric field (2)

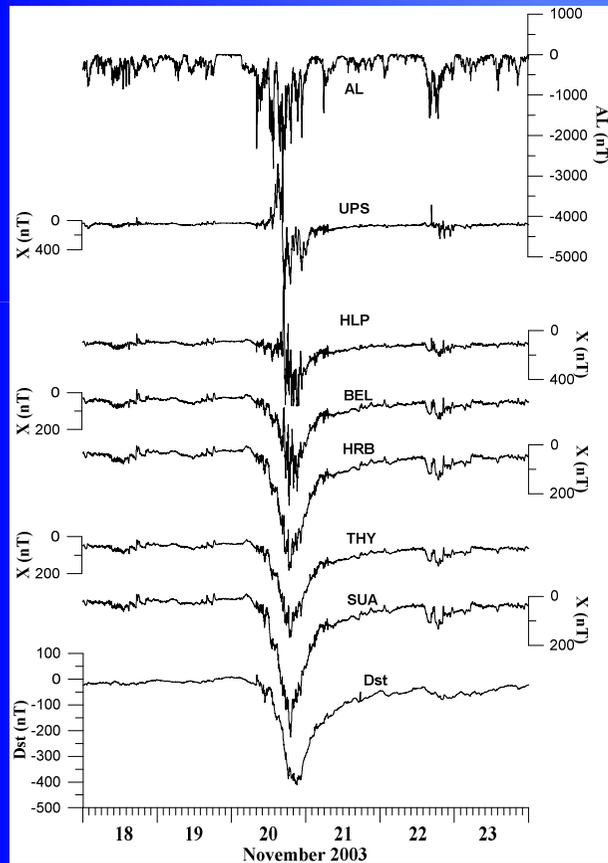
Conductivity model of the European GIC project



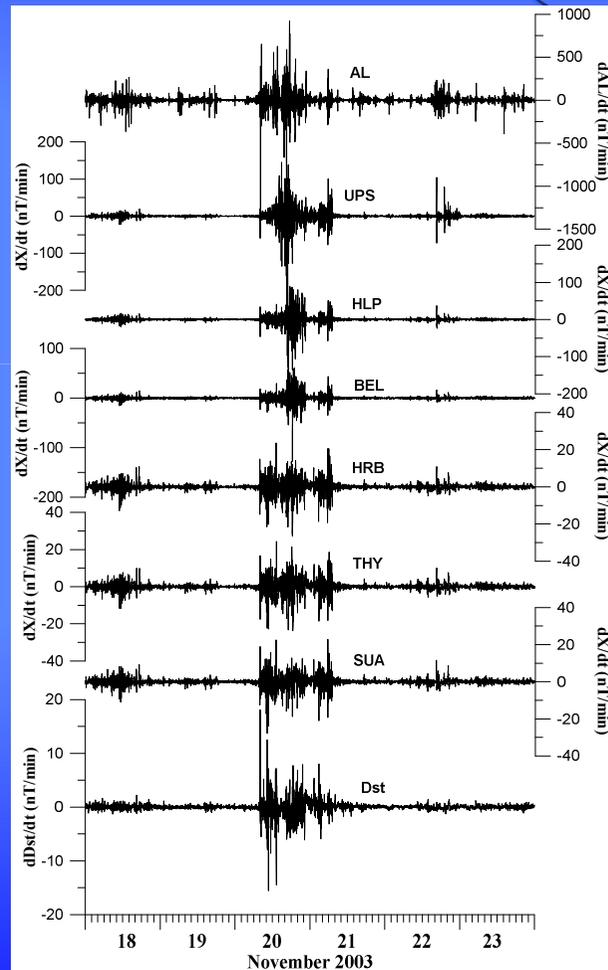
Surface geoelectric field (3)

November 2003 storm
~105°E

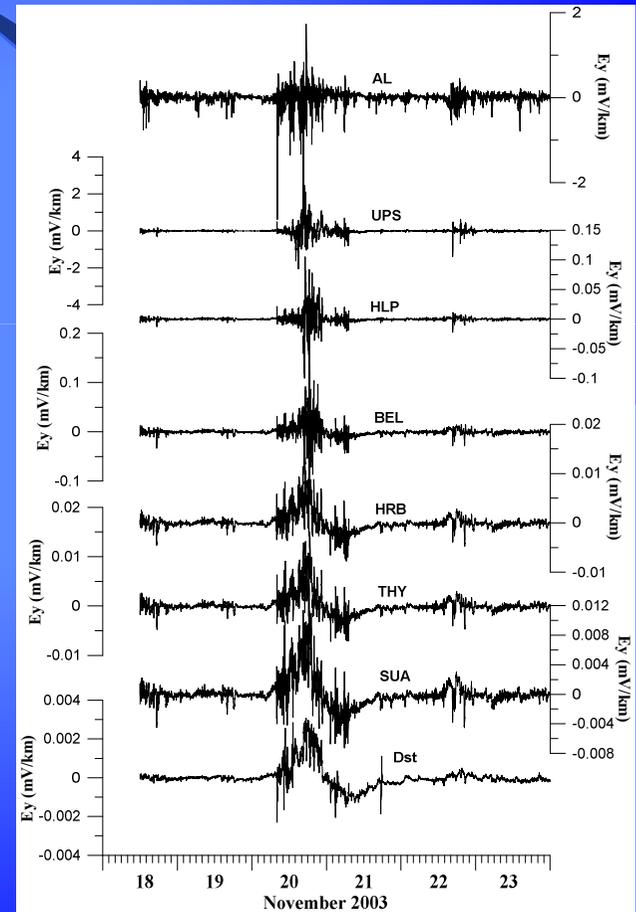
X (nT)



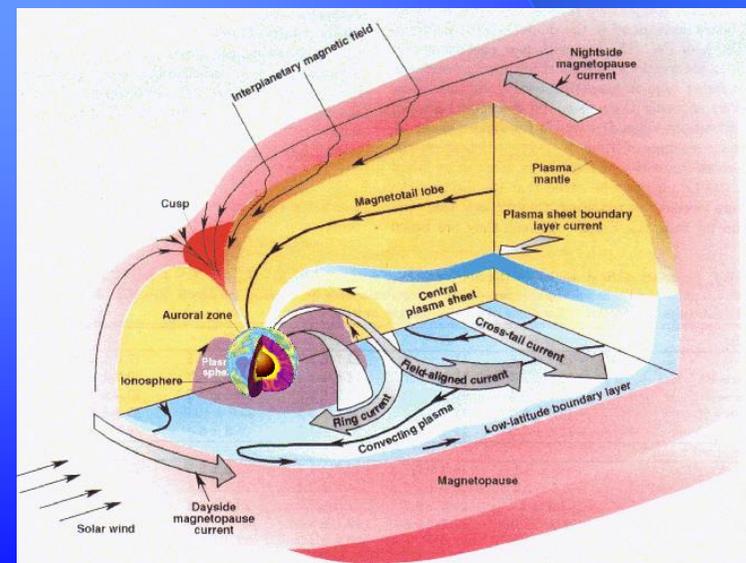
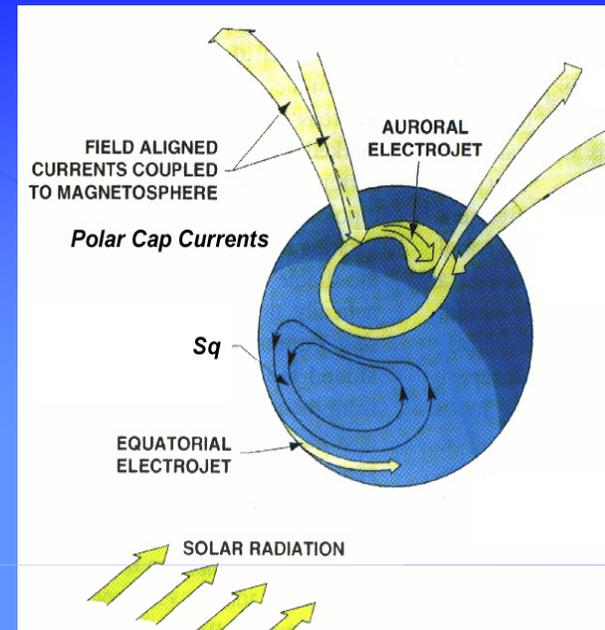
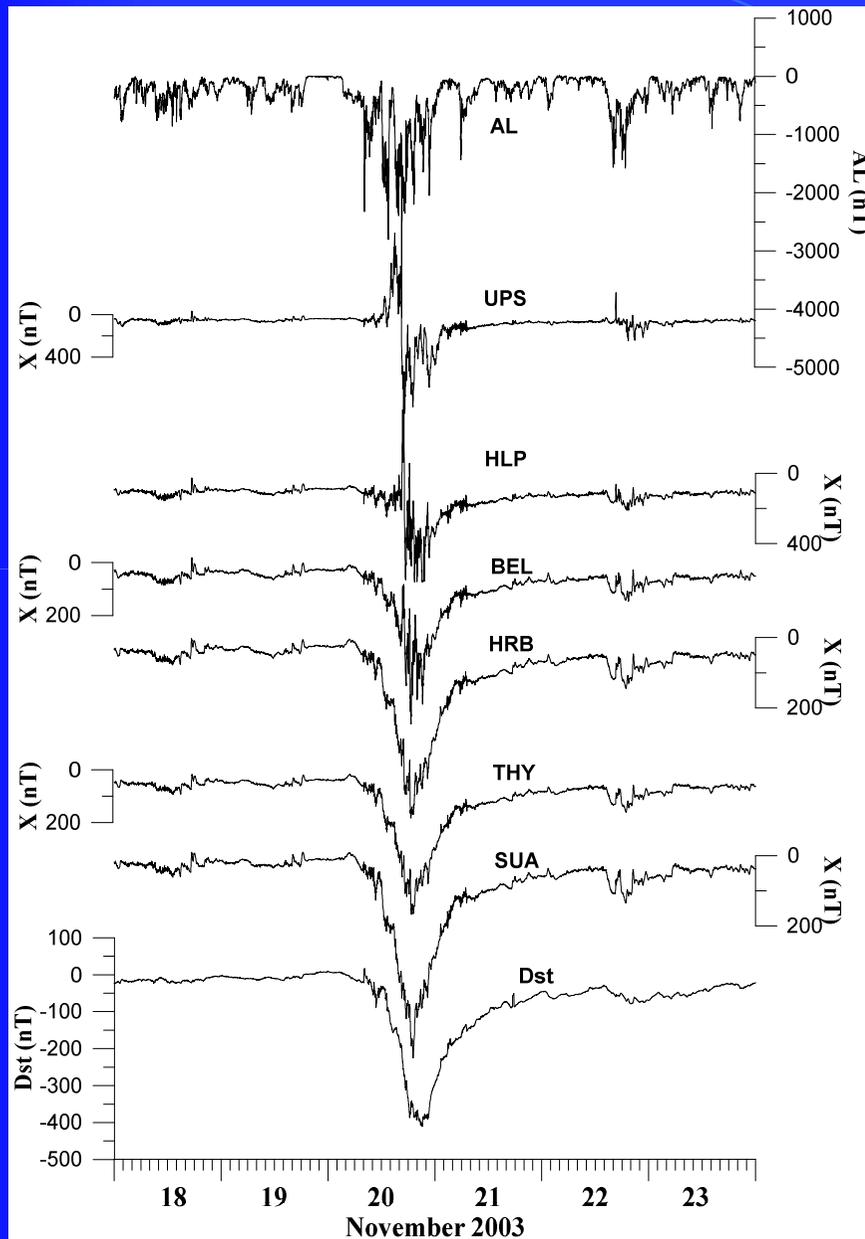
dX/dt (nT/min)



E_y (mV/km)

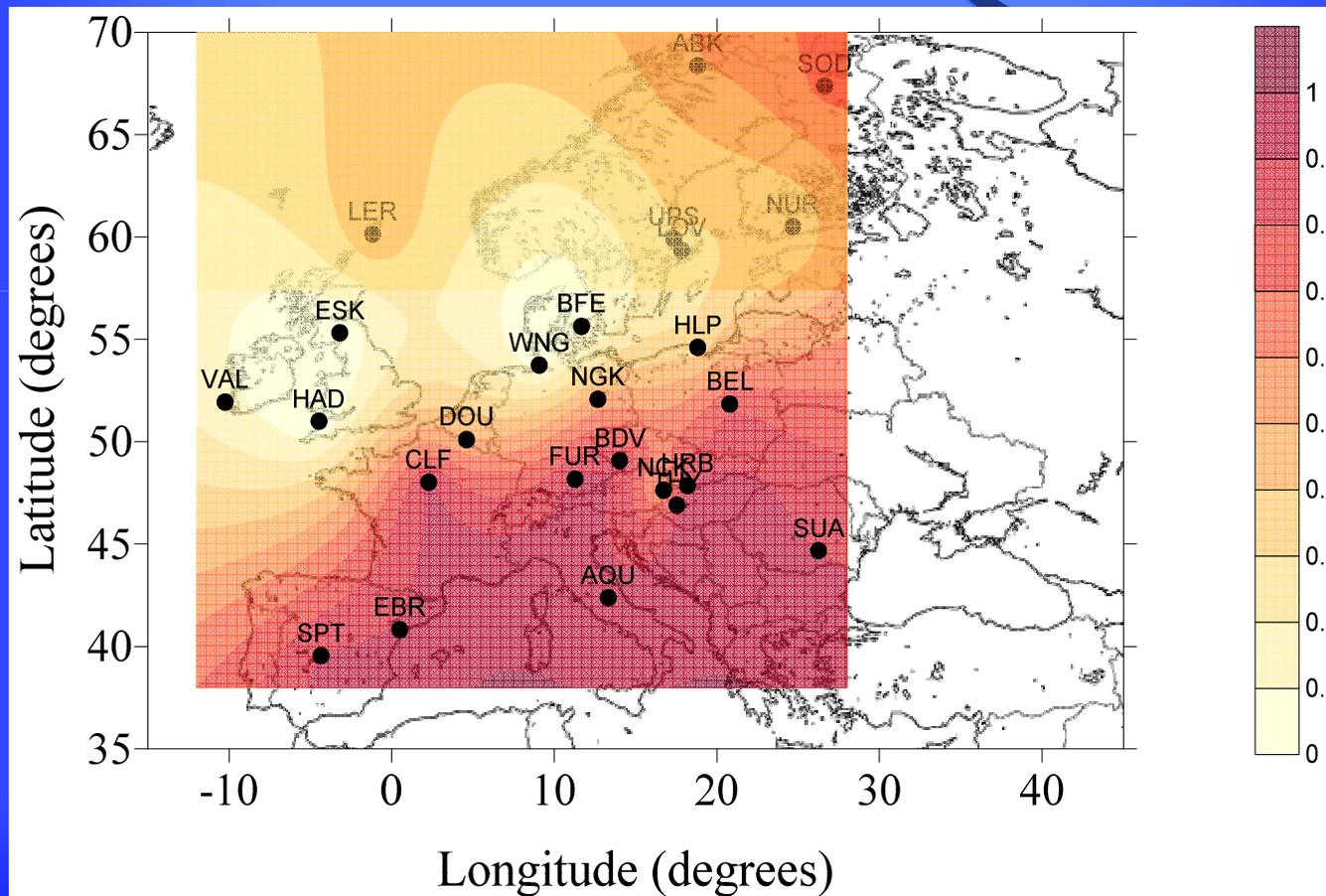


Sources of geomagnetic disturbance (1)

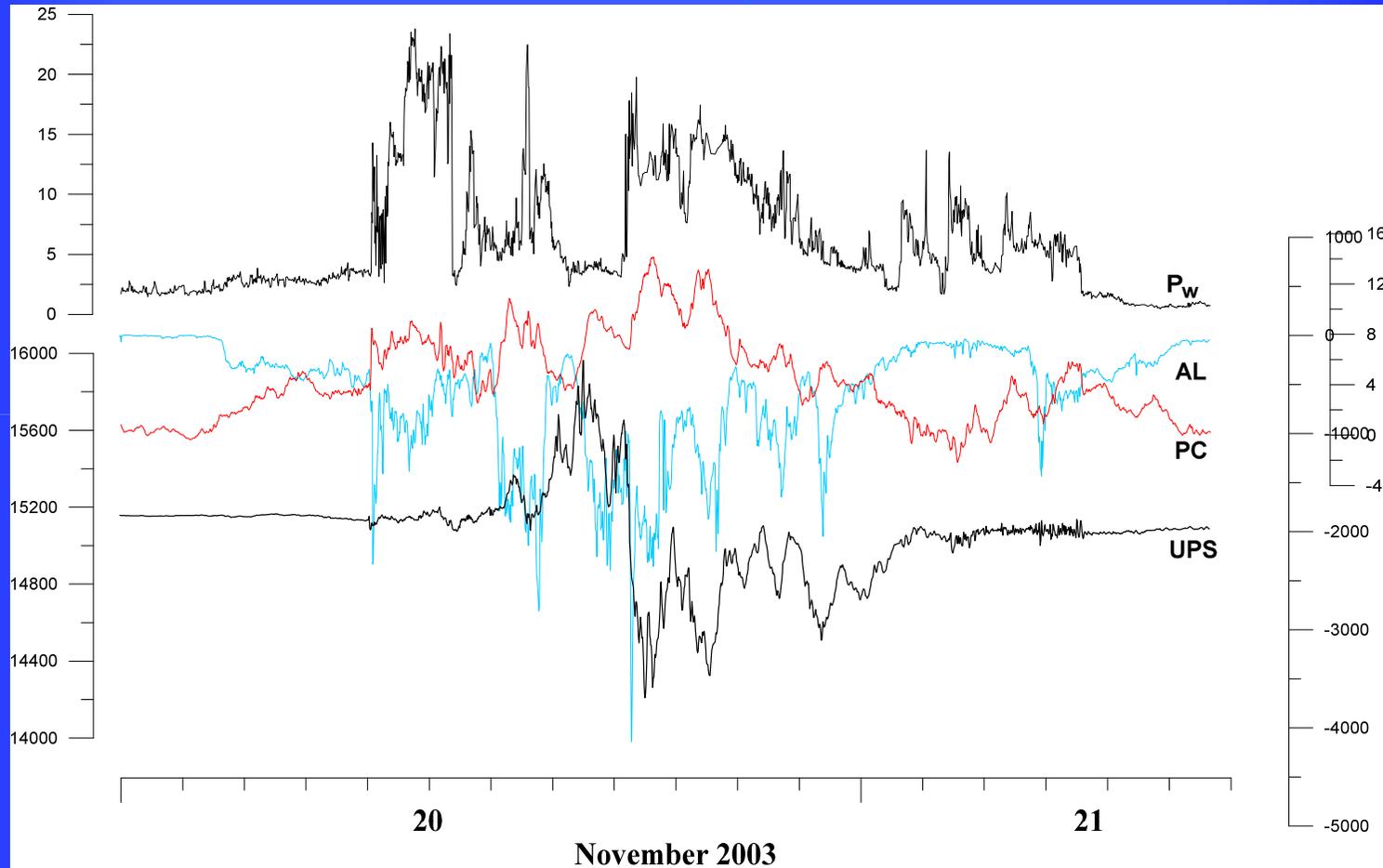


Sources of geomagnetic disturbance (2)

Ring current contribution

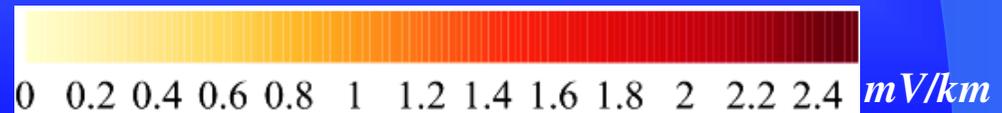
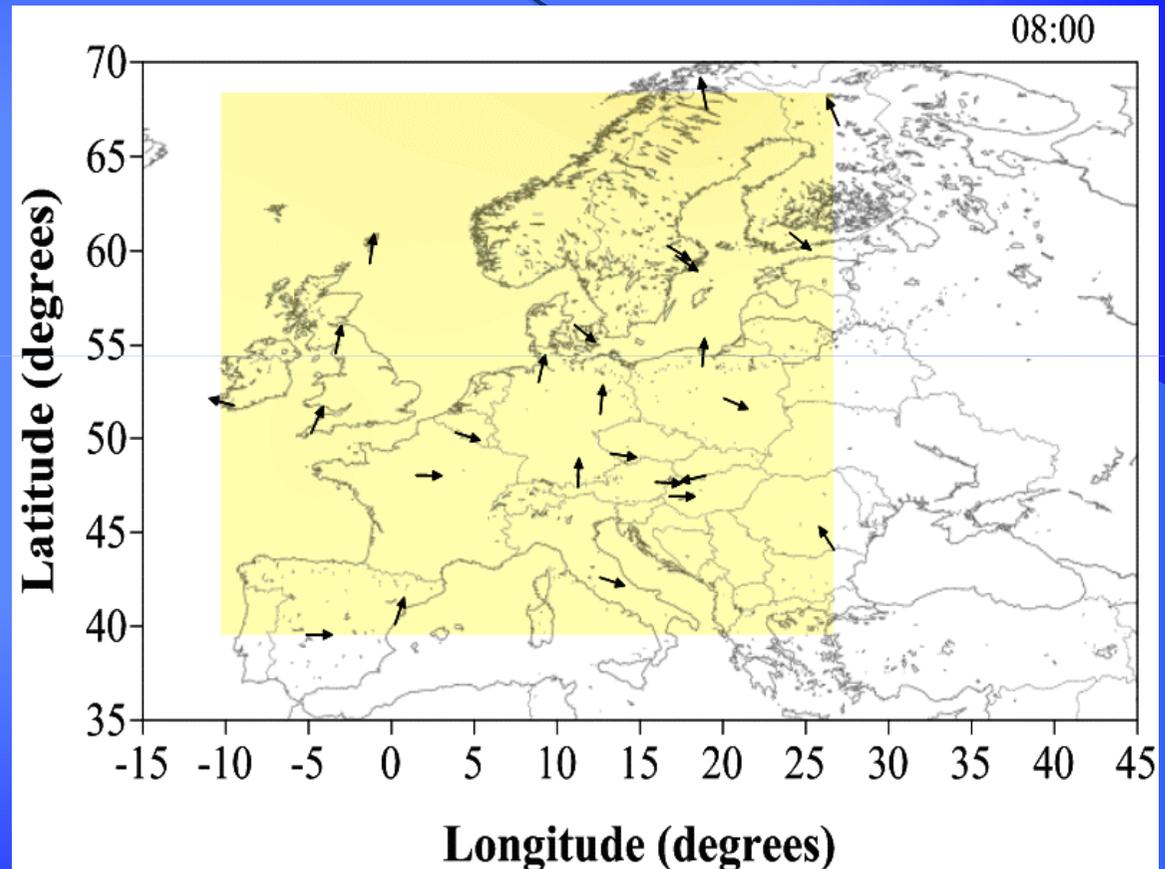
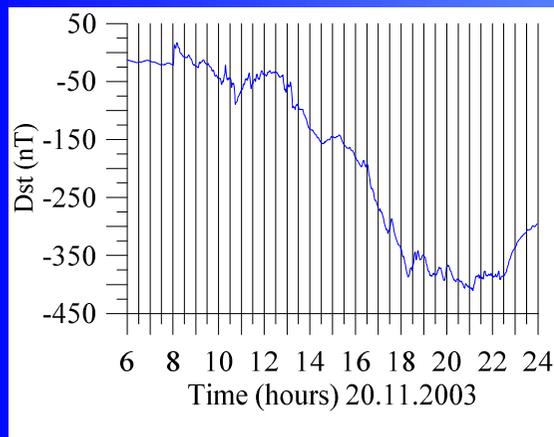
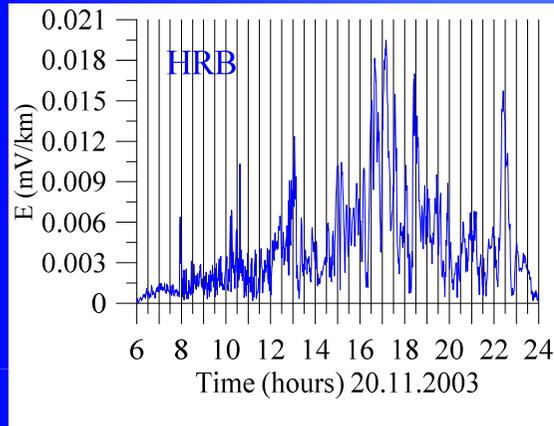


Sources of geomagnetic disturbance (3)



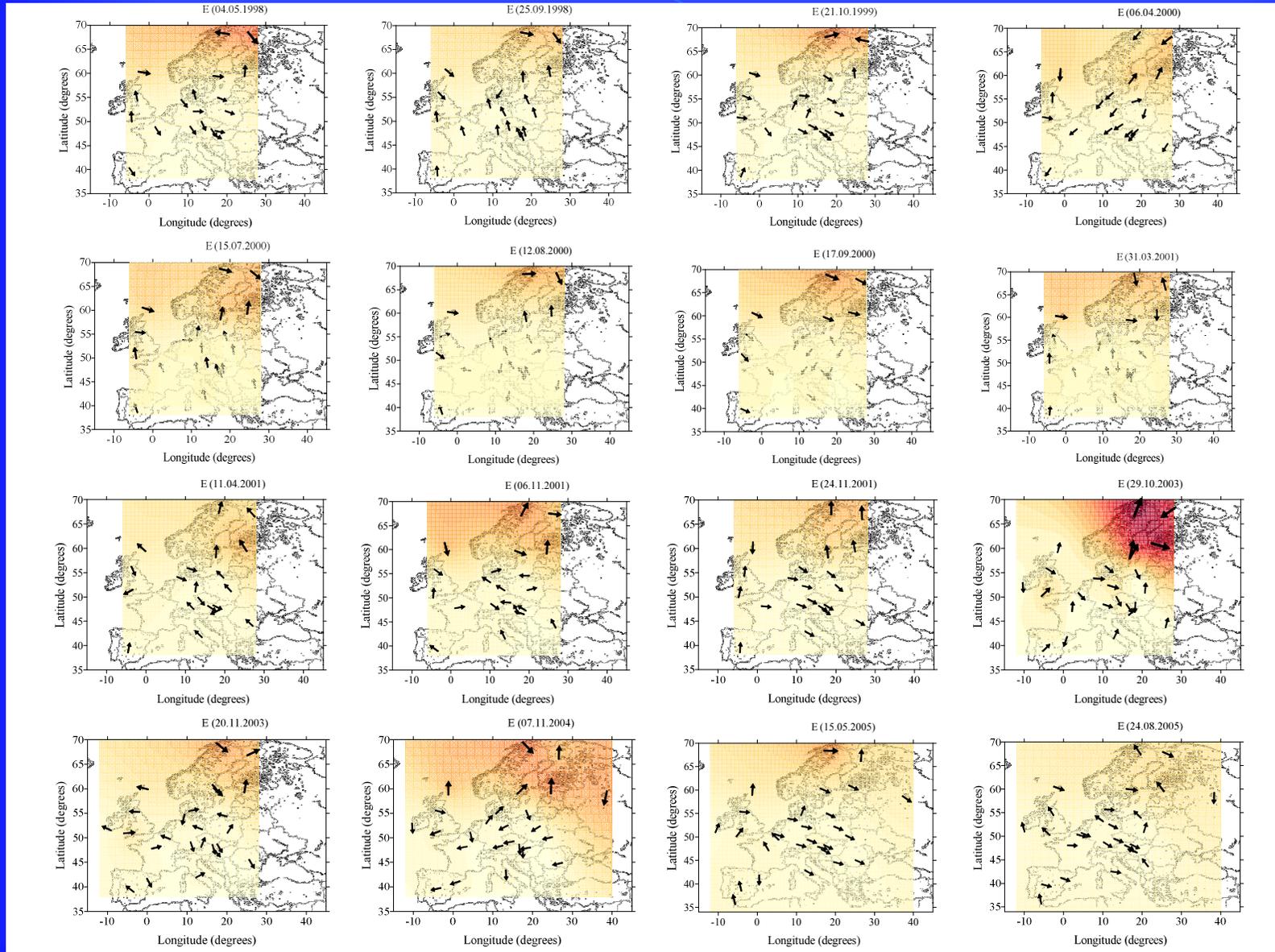
Geoelectric field evolution

Initial & main phase – November 2003 storm

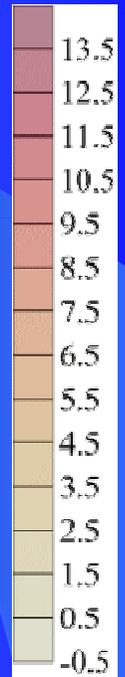


Toward GIC hazard assessment

E_{max} maps



mV/km



Conclusions

- *the disturbance in X is 2-3 times larger at northern latitudes than at mid&southern latitudes;*
- *during the geomagnetic storm, effects of auroral electrojets superimpose at all latitudes on the disturbance created by the magnetospheric ring current;*
- *the amplitude of the geoelectric field produced by magnetic variations is of the order of hundredths of mV/km in case of SUA (45°N), and of 1-2 mV/km in case of UPS (60°N);*
- *the maximum E value is not reached at the same moment at all observatories and its orientation depends on that moment of the storm development;*
- *future work: look at local effects and explore the role of magnetopause currents;*
- *the present approach concerns only the geophysical problem of GIC hazard. Engineering solutions are the next step.*