Analysis of the Geomagnetic Activity during the SC 24 Maximum Phase

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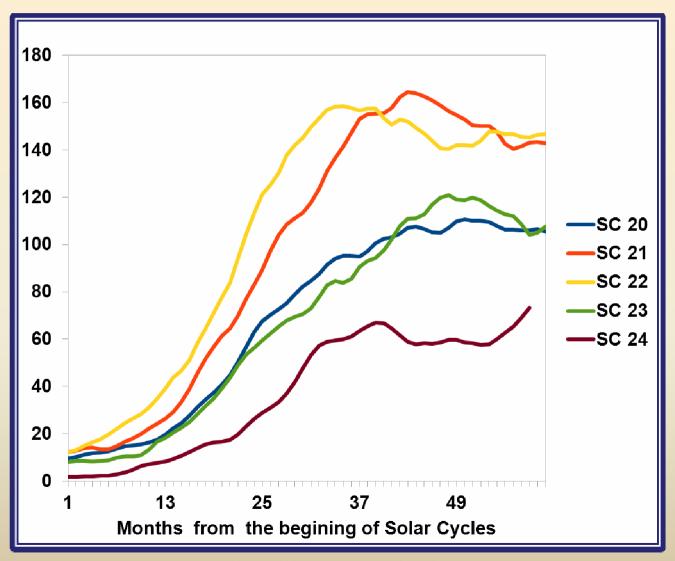
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Outline

- Introduction
 - Solar cycle 24 maximum phase;
 - Geomagnetic variability indices (Kp, aa, Ap ...), storms;
 - HSS and HSS_GS Catalogues for SC 23
- HSS Catalogue for SC 24 (2009 2013):
 - Data sources, HSS parameters;
 - Statistics
- SC 24 maximum phase:
 - Comparative analysis of R, aa, Ap and Σ I;
 - Event analysis: March 2012 and March 2013
- Future work

Solar Cycles nos. 20 – 24 (60 months)



SC 24 maximum phase



July 2011 – October 2013 (unfinished)

R(max.1) = 66.9February 2012

Geomagnetic Storms

Geomagnetic storm classification:

- $-50 < Dst \le -30 \rightarrow minor (small) storm (tipically substorm);$
- $-100 < Dst \le -50 \rightarrow moderate storm$;
- $-150 < Dst \le -100 \rightarrow intense (strong) storm;$
- - Dst \leq -150 \rightarrow severe storm.

Three distinct phases:

- •Initial phase;
- Main phase lasting to several hours —
 Dst can reach negative values of hundreds of nT;
- •Recovery phase lasting from tens of hours to a week Dst gradually returns to the normal level.

http://www.spacescience.ro/new1/Pagini finale/GS HSS Catalogue.htm.

HSS definition

HSS – a large increase in the SW velocity lasting by several days

- •Intriligator (1973) HSS as a stream having a rapidly rising increase in solar wind speed and a peak velocity \geq 450 km/s;
- •Bame et al. (1976) and Gosling et al. (1976) define a HSS as an observed variation of solar wind speed characterized by an increase of at least 150 km/s within a 5-day interval;
- •Broussard et al. (1977) define a HSS as wind period in which the solar wind speed is \geq 500 km/s averaged over a day

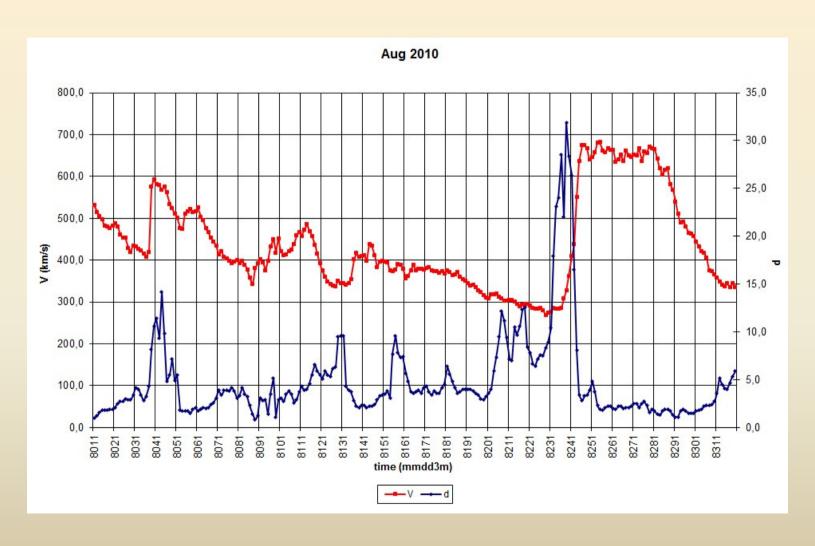
HSS Parameters

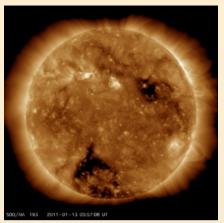
- HSS duration d (days);
- HSS initial velocity the smallest 3-hr velocity mean value for a given day, V_0 ;
- HSS maximum velocity V_{max} (km/sec);
- HSS velocity gradient $\Delta V_{max} = V_{max} V_o$;
- HSS importance (or intensity) $I = \Delta V_{max} \times d$

(ΣI could be used, calculated per BRs, per months, or per years, as well as, the weighted values : $I_p = \sum I / \sum d$)

All these parameters characterize in some manner the energy of the events; the <u>HSS Importance</u> is considered to be the best one because it includes the velocity gradient as well as the duration of the event.

August 2010

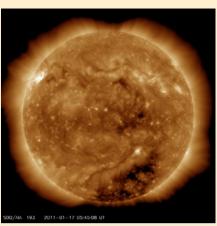


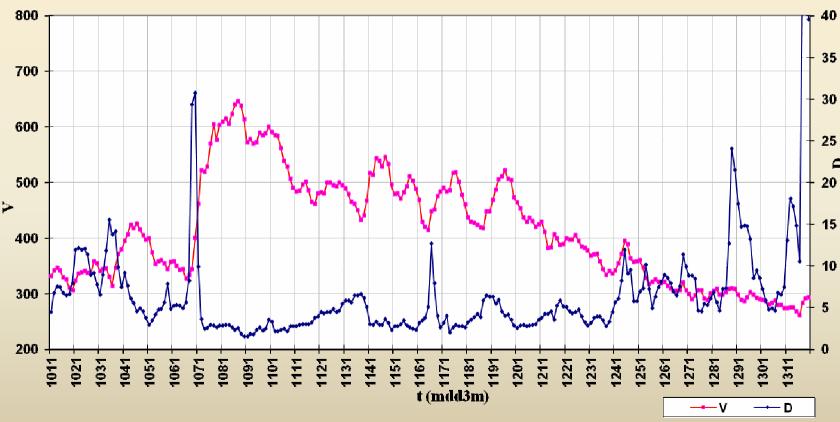


January 2011

Left: 13 January, Right:17 January

January 2011

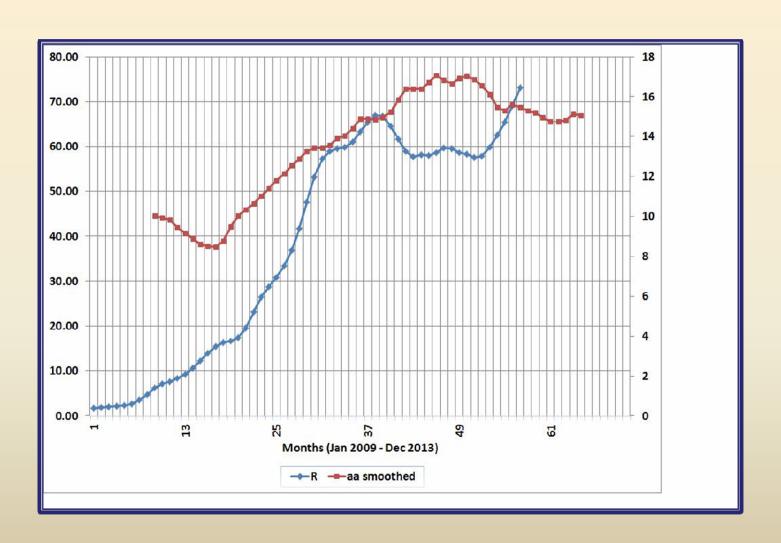




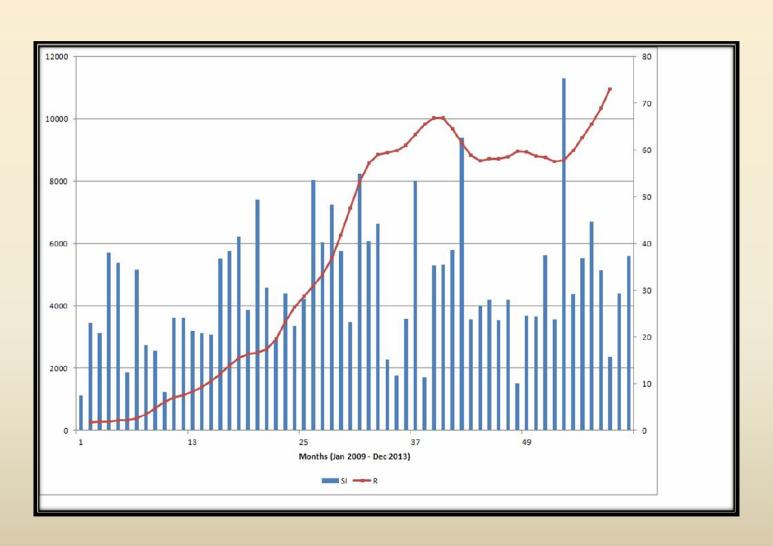
HSSs Statistics during Solar Cycles 23 and 24 (60 months)

Year	SC 23					SC 24					
rear	1996	1997	1998	1999	2000	2009	2010	2011	2012	2013	
all_HSSs	45	35	41	46	48	34	34	44	39	36	
CH_HSSs	43	24	26	36	25	34	28	27	31	29	
FG_HSSs	2	11	15	10	23	0	6	17	8	7	
HSSs/month	3.8	2.9	3.4	3.8	4.0	2.8	2.8	3.7	3.2	3	
Vmax>600 km/s	13	4	14	20	19	1	8	13	12	7	
Vmax>700 km/s	1	0	3	5	8	0	2	2	4	2	
∆V>400 km/s	0	0	2	2	5	0	2	0	0	3	

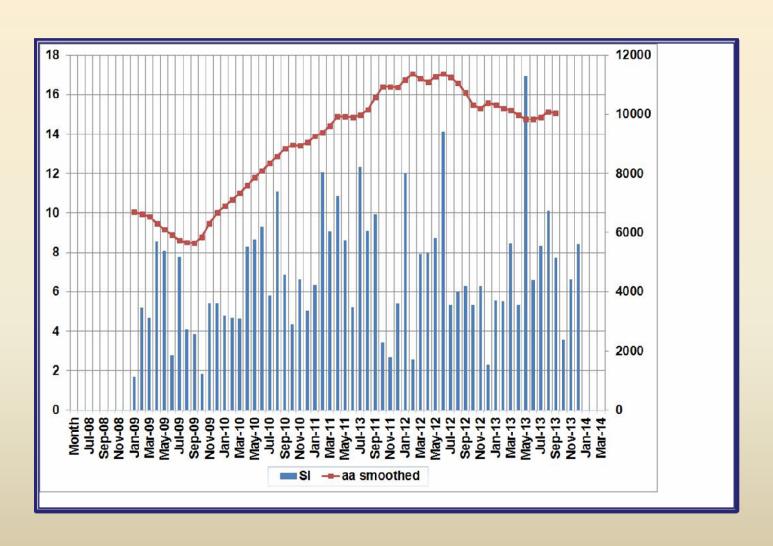
R and aa, smoothed monthly values

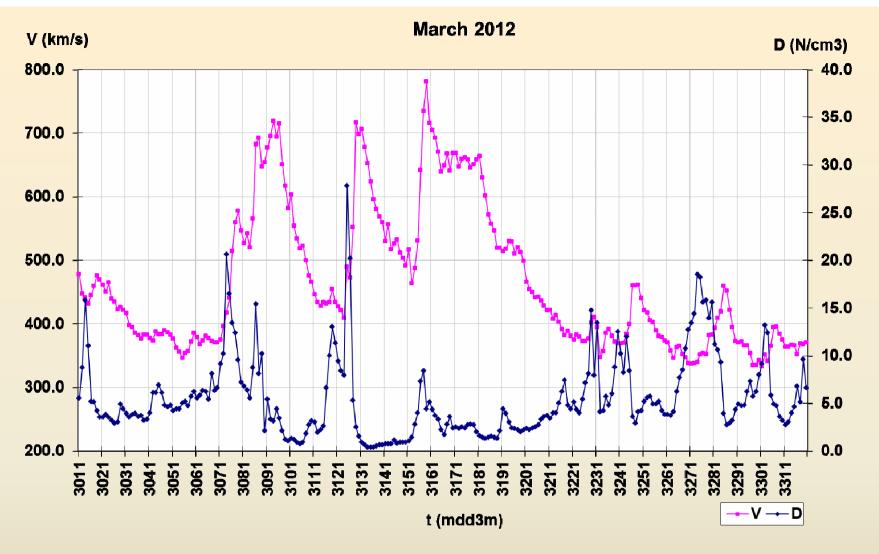


R and **∠1**, 2009 - 2013



2 and **aa**, 2009 - 2013

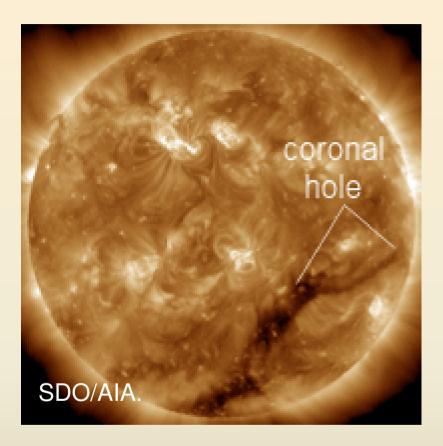


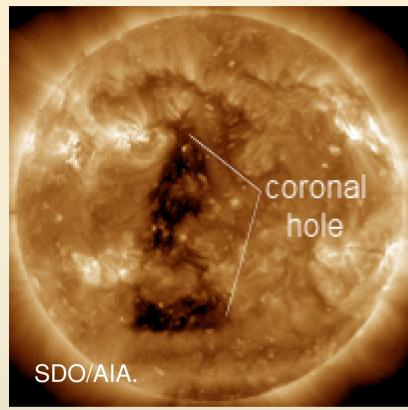


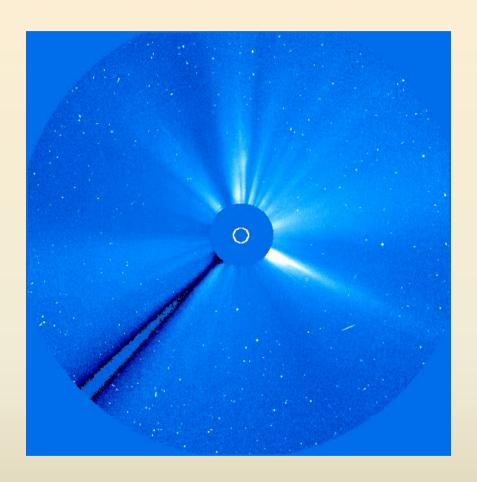
Υ	М	D	3-h	V0	V1	VMax	Dur	ΔV1	ΔVΜ	Source	I
2012	3	12	3	408.7	716.3	716.3	2.9	307.7	307.7	CH, ICME	892.33
2012	3	15	2	463.3	781	781	5	317.7	317.7	CH, ICME	1588.5

8 March 2012

13 March 2012





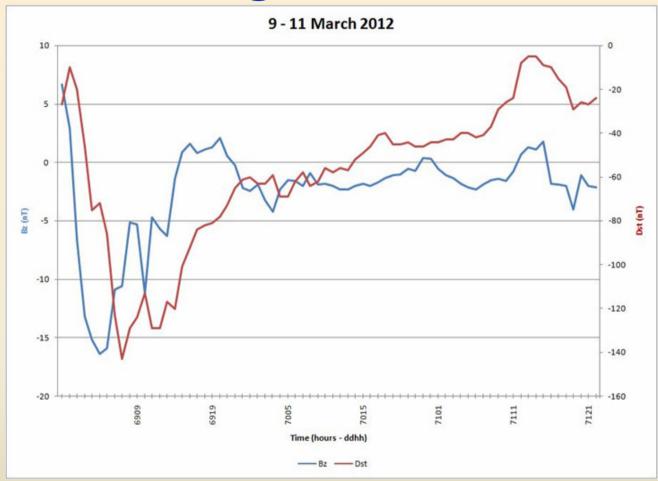


An ICME was recorded on March 9, 05:00 – March 11, 07:00 listed in the Richardson and Cane catalogue, correlated with a Halo CME from March 7, 00:24.

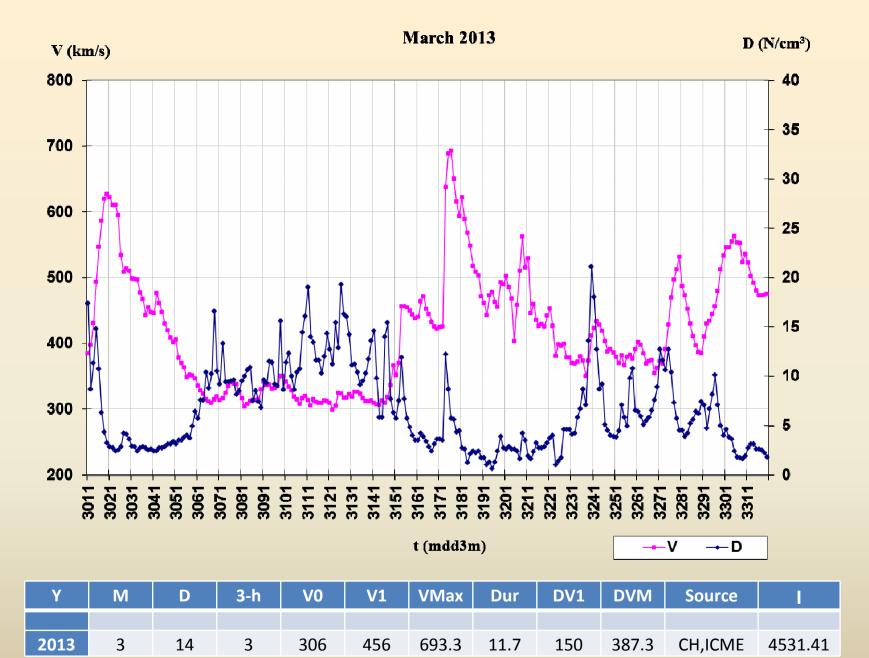
ICME shows evidence of a rotation in field direction, but lacks some other characteristics of a magnetic cloud.

Vm (ICME) = 550 km/sV max = 890 km/s

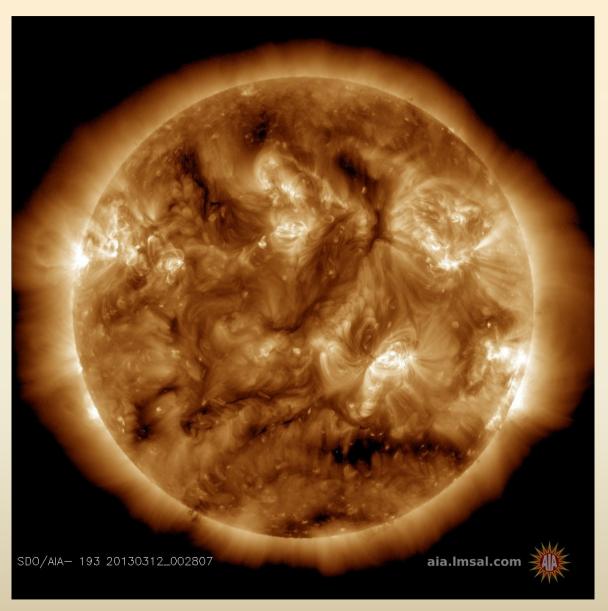
Geomagnetic storm



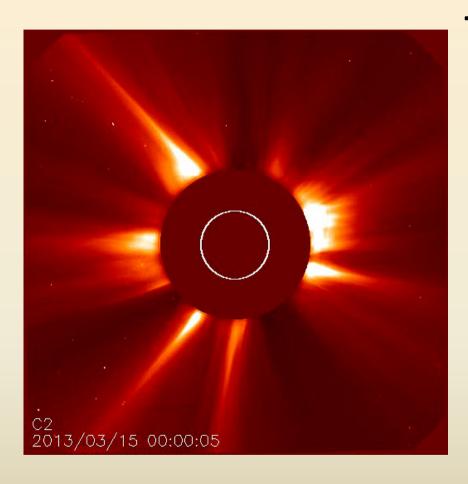
An intense geomagnetic storm with Sudden Commencement (minimum Dst = -131nT; provisional value) was registered on March 9, 2012.



12 March 2013



ICME; CME



Two ICMEs were recorded:

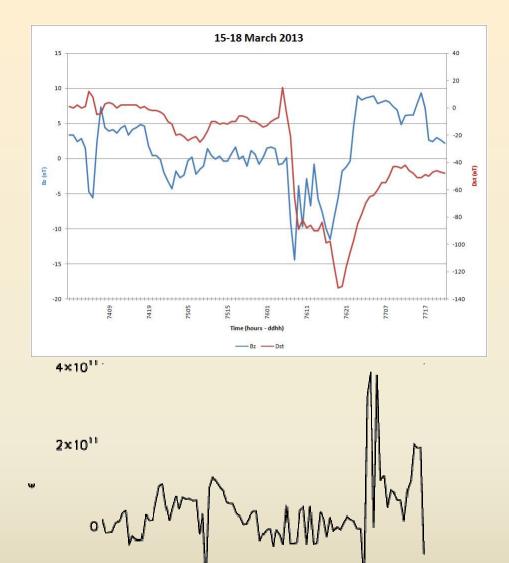
- March17, 15:00 -19, 17:00;

Vm = 520 km/s, V Max = 720km/s, a magnetic cloud reported), correlated with a Halo CME from March 15, 07:12.

- March 20, 18:00-22:00

Vm = 520 km/s, V Max = 640km/s, ICME shows evidence of a rotation in field direction, but lacks some other characteristics of a magnetic cloud)

26 - 30 May 2014



2013-Mor-16 Date

March 17: an intense geomagnetic storm with sudden commencement was registered (minimum Dst = -132)

2013-Mor-14

 -2×10^{11} .

2013-Mor-18

Future work

- HSS and HSS_GS Catalogues for SC 24
 - –Solar sources → CH, eruptive event (CME, F, SDF, AP);
 - Interplanetary magnetic field (B, Bz);
 - —I and Ip dynamics;
 - Geomagnetic effects
- Energy transfer from SW to the terrestrial magnetosphere

Thank You for Your kind attention!

HSS Catalogues – SC23 (1996 – 2008); SC24 (2009-2013)

HSS_Catalogue.html

COLUMN SIGNIFICATION:

- •Columns 1–6: the start data of the streams by: **Y**–year; **M**–month; **D**–day (calendar data); **3-H**–3-hr interval of the start day; **BR**–Bartels Rotation; **D-BR** day of Bartels Rotation, corresponding to the calendar day of the stream start;
- •Column 7: **V0**–minimum (pre-stream) velocity;
- •Column 8: V1–maximum velocity in the second day of the stream;
- •Column 9: **dt1**–time interval between **V0** and **V1** (in number of 3-hr intervals);
- •Column 10: VMax-maximum velocity of the stream;
- •Column 11: **Dur** duration of the stream, in days;
- •Column 12: $\Delta V1 = V1 V0$ —gradient of the velocity;
- •Column 13: $\Delta VM = VMax-V0$ —maximum gradient of the plasma velocity;
- •Column 14: **IMF**—the dominant polarity of the IMF for the duration of stream (+/– or –/+ means a magnetic sectorial border);
- •Column 15: **Source** solar source of the stream: **CH** –coronal hole; **F** solar flares or other eruptive phenomena.

HSS Catalogue

Maris G. and Maris O., 2012, High speed streams in the solar wind during the 23rd solar cycle, in: Advances in Solar and Solar-Terrestrial Physics, Editors: Maris G. and Demetrescu, C, Chapter 7, Published by Research Signpost, T.C. 37/661(2), Fort P.O., Trivandrum-695 023, Kerala, India, ISBN <u>978-81-308-0483-5</u>, pp. 97-134

www.spacescience.ro/new1/HSS_Catalogue.html